

**Peekskill Meteorite**

The car and the meteorite

Now in the Smithsonian!



A meteorite the size of a car streaks across the sky, headed for Canada, but doesn't strike the ground. It "skips" off the atmosphere.



**Sometimes they hit houses**

Wethersfield, Connecticut, November 8, 1982.

6-pound meteorite crashed through roof.



**Some meteorites are big**

15-ton meteorite found in the Willamette Valley, Oregon

Known to native Americans, but "discovered" in 1902. Now in the Hayden Planetarium, New York City.



1906 photo



This is an iron meteorite

**Many meteorites have been recovered in Antarctica**

Meteorites are preserved in the Antarctic ice cap.

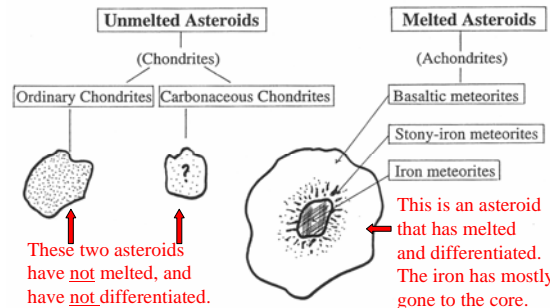
They are concentrated in certain areas by the motion of the ice.

Meteorites are easier to find in the ice because there are very few "native" rocks.

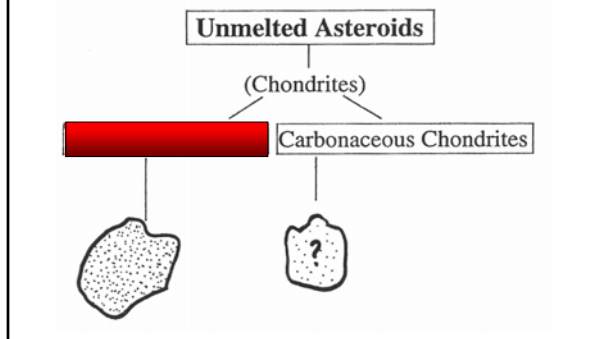
A few thousand have been found so far.



**Melted and unmelted meteorites come from different kinds of asteroids**



Let's consider the unmelted ones first



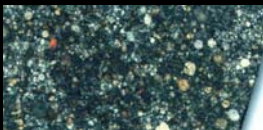
Chondrite meteorites

- Chondrites are meteorites that contain **chondrules**
- **Chondrules** are little BB-size blobs of minerals and mineral glass that were once melted and quickly solidified. They look like gravel.



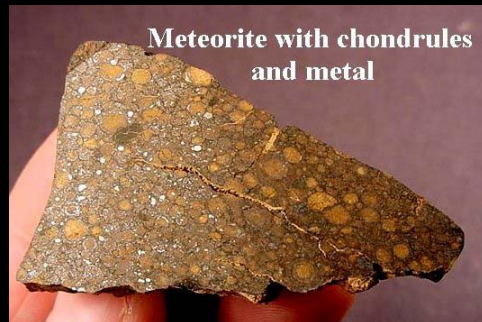
Chondrules and the early Solar System

- The chondrules in the **ordinary chondrite** meteorites formed very early in the history of the Solar System, before the Sun and planets had completely formed.
- We don't know the details of their origin.
- Ordinary chondrite meteorites come from asteroids that consist of this ancient material.



Section of an ordinary chondrite meteorite. The chondrules are about 1 to 4 mm in size.

- **Ordinary chondrites** contain metal (iron and nickel). Note the tiny shiny flakes in this meteorite.

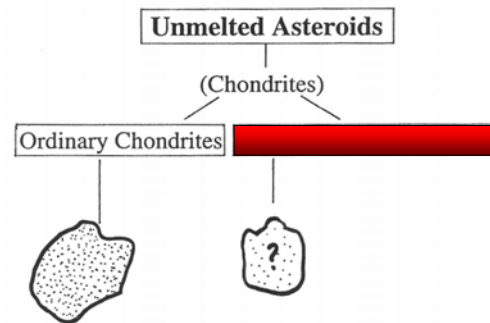


Very important !

- The **ordinary chondrites** have the overall composition (metal + silicate minerals) of the Earth and the other terrestrial planets.
- If we put the Earth in a giant blender and ground it all up (*core, mantle, and crust*) into small pieces, it would have the overall composition of an ordinary chondrite.




Now the Carbonaceous Chondrites




### Carbonaceous Chondrites

- These meteorites contain a significant amount of *carbon*, in addition to the chondrules and metal.
- Much of the carbon is in the form of complex organic molecules, including many amino acids.
- Two of the most famous carbonaceous meteorites are Allende and Murchison; both fell in 1969.




Allende  
(Mexico)




Murchison  
(Australia)

### Carbonaceous Chondrites

- These meteorites may have brought water to the early Earth
- They also brought complex organic materials that *may* have aided in the origin of life on Earth.



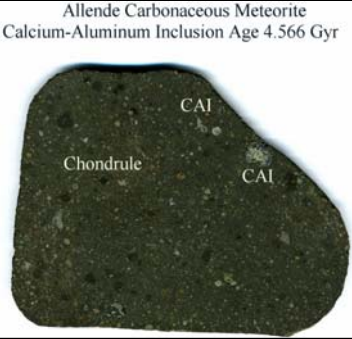
Allende  
(Mexico)



Murchison  
(Australia)

The carbonaceous meteorites are extremely old. The calcium-aluminum inclusions (CAIs) in them are the oldest known matter in the Solar System.

Allende Carbonaceous Meteorite  
Calcium-Aluminum Inclusion Age 4.566 Gyr



The CAIs solidified 4.566 billion years Ago.


We determine the age of the Solar System from the ages of the meteorites

### Again, melted and unmelted meteorites come from different kinds of asteroids


Unmelted Asteroids

(Chondrites)

Ordinary Chondrites



Carbonaceous Chondrites



These two asteroids have not melted, and have not differentiated.

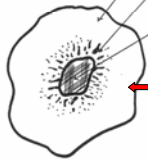
Melted Asteroids

(Achondrites)

Basaltic meteorites

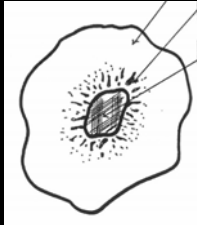
Stony-iron meteorites

Iron meteorites



This is an asteroid that has melted and differentiated. The iron has mostly gone to the core.

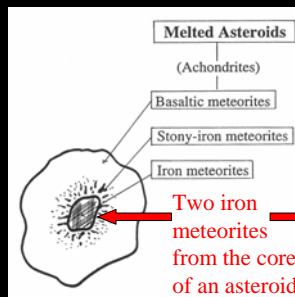
This sketch shows a cut through an asteroid that has melted and *mostly differentiated*. The iron and other metals were sinking to form a core, and the lighter silicate minerals were floating to the top. The asteroid was small, and solidified before differentiation was entirely complete.





The *outer layer* is made of silicate rocks.  
The *core* is made of iron + nickel.  
The *intermediate zone* is a mix of metal and silicate minerals.

We can call this a *differentiated asteroid*

When *differentiated* asteroids are shattered, they produce at *least three kinds* of meteorites.



Two iron meteorites from the core of an asteroid

When differentiated asteroids are shattered, they produce at least three kinds of meteorites.

Melted Asteroids  
(Achondrites)

- Basaltic meteorites
- Stony-iron meteorites
- Iron meteorites

A stony-iron meteorite from the region near the core. Note the green olivine crystals.

When differentiated asteroids are shattered, they produce at least three kinds of meteorites.

Melted Asteroids  
(Achondrites)

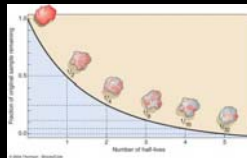
- Basaltic meteorites
- Stony-iron meteorites
- Iron meteorites

A basaltic meteorite from the volcanic crust of a differentiated asteroid

### The Importance of Meteorites

- They are the oldest material that we have.
- Their ages are consistently 4.45 to 4.57 billion years.
- Meteorite ages are determined from the radioactive atoms in them.
- You have seen these two figures before ...

Parent	Daughter	Half-Life (billion yr)
Samarium-147*	Neodymium-143	106
Rubidium-87	Strontium-87	48.8
Thorium-232	Lead-208	14.0
Uranium-238	Lead-206	4.47
Potassium-40	Argon-40	1.31



### A final critical point ...

- The age of the formation of the Solar System (that is, the condensation of solid material from gas and dust) is based on the ages of the meteorites.
- This is where we get the figure **4.6 billion years** for the age of the Earth, the other planets, and the Sun
- It is based on hundreds of independent age measurements on hundreds of different meteorites.